On the design of density-functional approximations for canonical and grand-canonical ensembles

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Optical and fundamental gaps in DFT

• The difference in energy between the *N*-electron first excited and ground states is referred to as the optical gap. It is expressed as follows,

$$\omega^N = E_1^N - E_0^N$$

- Note that the optical gap describes a neutral excitation.
- On the other hand, the **fundamental gap** is a ground-state quantity.
- It describes the removal and addition of an electron to the system.
- It is calculated from the electron affinity [EA] A^N and the ionization potential [IP] I^N :

$$E_{g}^{N} = \underbrace{\left(E_{0}^{N-1} - E_{0}^{N}\right)}_{I^{N}} - \underbrace{\left(E_{0}^{N} - E_{0}^{N+1}\right)}_{A^{N} = I^{N+1}}$$

 Interestingly, for non-interacting electrons, the two gaps boil down to the same quantity, namely the HOMO-LUMO gap of the N-electron system:

$$\omega^{\rm N} = {\cal E}_{\rm g}^{\rm N} = \varepsilon_{\rm L}^{\rm N} - \varepsilon_{\rm H}^{\rm N}.$$

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Limitations of the one-electron picture



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DFT for a fixed (integral) number N of electrons

 In quantum mechanics, the electronic repulsion energy is expressed (in atomic units) as follows,

$$W_{\mathrm{ee}} = \frac{N(N-1)}{2} \int_{\mathbb{R}^3} \mathrm{d}\mathbf{r}_1 \int_{\mathbb{R}^3} \mathrm{d}\mathbf{r}_2 \int_{\mathbb{R}^3} \mathrm{d}\mathbf{r}_3 \dots \int_{\mathbb{R}^3} \mathrm{d}\mathbf{r}_N \ \frac{\Psi^2(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, \dots, \mathbf{r}_N)}{|\mathbf{r}_1 - \mathbf{r}_2|},$$

where $\Psi(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, \dots, \mathbf{r}_N)$ is the *N*-electron wavefunction.

- In standard Kohn–Sham (KS) density-functional theory (DFT), the latter energy can be determined, in principle exactly, from the electron density $n(\mathbf{r})$, which is mathematically a much simpler quantity than the wavefunction.
- For that purpose, the so-called Hartree-exchange-correlation (Hxc) density functional has been introduced,

$$W_{\rm ee} \to E_{\rm Hxc}[n],$$

where the density is determined exactly from the KS orbitals as follows,

$$n(\mathbf{r}) = \sum_{i=1}^{N} |\varphi_i(\mathbf{r})|^2.$$

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• Self-consistent equations fulfilled by the KS orbitals:

$$-\frac{1}{2}\nabla^{2}\varphi_{i}(\mathbf{r})+\left[\mathbf{v}_{\mathrm{ext}}(\mathbf{r})+\frac{\delta \mathbf{E}_{\mathrm{Hxc}}[\mathbf{n}]}{\delta \mathbf{n}(\mathbf{r})}\right]\times\varphi_{i}(\mathbf{r})=\varepsilon_{i}\varphi_{i}(\mathbf{r}),$$

where $v_{\text{ext}}(\mathbf{r})$ is any external local (multiplicative) interaction potential energy (the nuclear-electron attraction potential for example) at position \mathbf{r} .

• The additional Hxc potential $\delta E_{\text{Hxc}}[n]/\delta n(\mathbf{r})$ ensures that the density of the true system is recovered, in principle exactly, from the KS orbitals.



Do the KS orbital energies have any physical meaning ?

- Let us denote $\varepsilon_{\rm H}$ and $\varepsilon_{\rm L}$ the Kohn–Sham HOMO and LUMO energies, respectively.
- In the following, we will focus on the (in-principle) exact calculation of the ionization potential (IP)

$$I_N = E_{N-1} - E_N$$

and the electron affinity (EA)

$$A_N = E_N - E_{N+1} = I_{N+1}$$

of the molecule within DFT.

- The energies E_{N-1} , E_N , and E_{N+1} of the cationic, *N*-electron and anionic molecule, respectively, should in principle be computed with different KS orbitals and energies.
- A simple but non trivial question: can we calculate both IP and EA from the KS energies and orbitals of the *N*-electron molecule ?
- In other words: $I_N \stackrel{?}{=} -\varepsilon_H$, $A_N \stackrel{?}{=} -\varepsilon_L$, and $I_N A_N \stackrel{?}{=} \varepsilon_L \varepsilon_H$.
- This is known as the fundamental gap problem in DFT.

Chemical potential in DFT

• For a fixed (integral) number N of electrons, shifting the local potential by a constant $-\mu$ has no impact on the KS orbitals and therefore no impact on the density:

$$\begin{aligned} -\frac{1}{2}\nabla^{2}\varphi_{i}(\mathbf{r}) + \left[\mathbf{v}_{\mathrm{ext}}(\mathbf{r}) + \frac{\delta E_{\mathrm{Hxc}}[n]}{\delta n(\mathbf{r})} - \mu\right] \times \varphi_{i}(\mathbf{r}) &= \varepsilon_{i}\varphi_{i}(\mathbf{r}) - \mu\varphi_{i}(\mathbf{r}) \\ &= \left(\varepsilon_{i} - \mu\right)\varphi_{i}(\mathbf{r}), \end{aligned}$$

• However, the total energy of the KS system is affected by such a shift:

$$\sum_{i=1}^{N} \varepsilon_i \quad \to \quad \sum_{i=1}^{N} \left(\varepsilon_i - \mu \right) = \left(\sum_{i=1}^{N} \varepsilon_i \right) - \mu N.$$

- As readily seen, the impact on the energy depends on the number of electrons.
- This is the reason why the constant μ , which is nothing but the chemical potential, plays a crucial role in grand canonical ensemble DFT (where N can vary).

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• The Hubbard atom is a model system consisting of a single (atomic) site that can contain up to 2 electrons:



• Starting from the one-electron atom case (N = 1) the IP and the EA read

$$I^1=E_0-E_1=-v_{\mathrm{ext}}$$
 and $A^1=E_1-E_2=-v_{\mathrm{ext}}-m{U}$, respectively,

so that the fundamental gap equals $F_g^1 = I^1 - A^1 = U$.

- In this simple (one-orbital atomic) model the HOMO-LUMO gap equals zero.
- Thus we conclude that, in general, the fundamental gap is *not* equal to the KS HOMO-LUMO gap.

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• If we now introduce the chemical potential μ , the energies are modified as follows,



• By varying μ we will be able to change the number of electrons N continuously^{*}.

- When μ < v_{ext}, the lowest energy is *E*₀ thus leading to N = 0.
 When v_{ext} < μ < v_{ext} + U, the lowest energy is *E*₁ thus leading to N = 1.
- When $\mu > v_{\text{ext}} + U$, the lowest energy is \mathcal{E}_2 thus leading to N = 2.

* Principles of DFT, lecture given by **Trygve Helgaker** at the GDR Correl mini-school on mathematics in electronic structure theory, Paris, January 2017.

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- What happens when $\mu = v_{\text{ext}}$?
- In the latter case, the quantum states $|0\rangle$ and $|1\rangle$ are strictly degenerate (${\cal E}_0={\cal E}_1=0).$
- Consequently, it is possible to mix them (and the energy will remain equal to zero) in order to obtain an N-electron quantum state,

$$\ket{\mathcal{N}} = \sqrt{1-\mathcal{N}} \ket{0} + \sqrt{\mathcal{N}} \ket{1}$$

where $0 < \mathcal{N} < 1$.

- The number of electrons N can vary continuously from 0 to 1 when $\mu = v_{ext}$.
- Similarly, when $\mu = v_{ext} + U$, the states $|1\rangle$ and $|2\rangle$ are strictly degenerate $(\mathcal{E}_1 = \mathcal{E}_2 = -U)$.
- In the latter case, $|\mathcal{N}\rangle = \sqrt{2-\mathcal{N}} |1\rangle + \sqrt{\mathcal{N}-1} |2\rangle$ where $1 < \mathcal{N} < 2$.

• As a result, the number of electrons can vary continuously from 1 to 2 when $\mu = v_{ext} + U$.

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- What happens to the energy (from which we remove the chemical potential contribution) when $\mu = v_{ext}$?
- For the mixed *N*-electron quantum state $|N\rangle = \sqrt{1 N} |0\rangle + \sqrt{N} |1\rangle$, the expectation value for the energy reads

$$E\left(\mathcal{N}
ight) = \left(\sqrt{1-\mathcal{N}}
ight)^2 imes E_0 + \left(\sqrt{\mathcal{N}}
ight)^2 imes E_1,$$

thus leading to $E(\mathcal{N}) = (1 - \mathcal{N})E_0 + \mathcal{N}E_1$ when $0 \leq \mathcal{N} \leq 1$.

• Similarly, when $\mu = v_{\mathrm{ext}} + U$, we have $1 \leq \mathcal{N} \leq$ 2 and

$$E(\mathcal{N}) = (2 - \mathcal{N})E_1 + (\mathcal{N} - 1)E_2$$

• Important conclusion: the energy varies linearly (piecewise) with respect to the number ${\cal N}$ of electrons.

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number \mathcal{N} of electrons

- It is therefore possible to calculate both IP and EA from a single DFT calculation by differentiating the energy with respect to N around N = 1.
- Note that $\mathcal{N} \to 1^-$ and $\mathcal{N} \to 1^+$ limits are *not* the same !
- Designing density functionals able to reproduce such a feature still remains a challenge. This is known as the derivative discontinuity problem.

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IP theorem

- Let us denote ν a *fractional number* of electrons in the range $N 1 < \nu < N$
- In the extension of Kohn–Sham DFT to open systems, the *piecewise linear** ν -electron

energy $E(\nu) = E_0^{N-1} - I^N \times (\nu - N + 1)$ is written as $E(\nu) = \min_{\{\varphi_i\}_i} \left\{ E\left[\nu, \{\varphi_i\}_i\right] \right\}$ with

$$\begin{split} E\Big[\mathbf{\nu},\{\varphi_i\}_i\Big] &= \sum_{i=1}^{N-1} \left\langle \varphi_i \middle| -\frac{\nabla^2}{2} + \mathbf{v}_{\rm ne}(\mathbf{r}) \times \middle| \varphi_i \right\rangle + \left(\mathbf{\nu} - N + 1\right) \left\langle \varphi_N \middle| -\frac{\nabla^2}{2} + \mathbf{v}_{\rm ne}(\mathbf{r}) \times \middle| \varphi_N \right\rangle \\ &+ E_{\rm Hxc} \left[n^{\nu} \right], \end{split}$$

where the density $n^{\nu}(\mathbf{r}) = \sum_{i=1}^{N-1} |\varphi_i(\mathbf{r})|^2 + (\nu - N + 1) |\varphi_N(\mathbf{r})|^2$ integrates to ν .

• The minimizing (*v*-dependent) orbitals fulfill the KS equations:

$$\left(-\frac{\nabla^2}{2}+v_{\rm ne}(\mathbf{r})+\frac{\delta E_{\rm Hxc}\left[n^{\nu}\right]}{\delta n(\mathbf{r})}\right)\varphi_i^{\nu}(\mathbf{r})=\varepsilon_i^{\nu}\varphi_i^{\nu}(\mathbf{r}).$$

* J. P. Perdew and M. Levy, Phys. Rev. Lett. 51, 1884 (1983).

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IP theorem

- Note that E_{Hxc} [n] is "truly" universal here as it applies to an arbitrary number of electrons (not only integral ones).
- *IP theorem:* $\left| \frac{\mathrm{d}E(\nu)}{\mathrm{d}\nu} = \varepsilon_N^{\nu} \right| \longrightarrow \varepsilon_N^{\nu} = -I^N \longrightarrow I^N = -\varepsilon_\mathrm{H}^N$
- Consequence of the IP theorem:

$$E_g^N = I^N - I^{N+1} = \varepsilon_{\mathbf{H}}^{N+1} - \varepsilon_{\mathbf{H}}^N.$$

- Note that, in principle, the LUMO of the *N*-electron system is *not* the HOMO of the (N + 1)-electron one.
- Therefore, there is no reason to believe that $\varepsilon_{\rm H}^{N+1}$ is equal to $\varepsilon_{\rm L}^{N}$.
- Let us stress that this statement *does not give much hint* about the way the difference between the latter orbital energies can be modelled in practical calculations.
- If we simply write $\varepsilon_{\rm H}^{N+1} = \varepsilon_{\rm L}^{N} + \Delta_{\rm xc}^{\rm fun.}$, then we recover the standard expression

$$E_g^N = \varepsilon_{\rm L}^N - \varepsilon_{\rm H}^N + \Delta_{\rm xc}^{\rm fun.}$$

Why $\Delta_{\rm xc}^{\rm fun.}$ is referred to as *derivative discontinuity* ?

• Let us now look at what happens when the number of electrons ν crosses an integer N.

• Three situations should be considered in the KS equations:

$$\left(-\frac{\nabla^2}{2}+v_{\rm ne}(\mathbf{r})+\frac{\delta E_{\rm Hxc}\left[n^{\nu}\right]}{\delta n(\mathbf{r})}\right)\varphi_i^{\nu}(\mathbf{r})=\varepsilon_i^{\nu}\varphi_i^{\nu}(\mathbf{r}).$$

- <u>Situation 1</u>: $\nu = N \eta$ where $\eta \to 0^+ \leftarrow KS$ potential unique.
- <u>Situation 2</u>: $\nu = N + \eta$ \leftarrow KS potential unique.
- <u>Situation 3</u>: $\nu = N$ \leftarrow KS potential unique up to a constant!

• The Hartree potential $\frac{\delta E_{\rm H}[n]}{\delta n(\mathbf{r})} = \int d\mathbf{r}' \frac{n(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|}$ is a *continuous* functional of the density.

• Therefore
$$\frac{\delta E_{\rm xc} \left[n^{N+\eta} \right]}{\delta n(\mathbf{r})} - \frac{\delta E_{\rm xc} \left[n^{N-\eta} \right]}{\delta n(\mathbf{r})} = \text{constant} = \varepsilon_{\rm H}^{N+1} - \varepsilon_{\rm L}^{N} = \Delta_{\rm xc}^{\rm fun.}.$$

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N-centered ensemble DFT



integral number M of electrons

B. Senjean and E. Fromager, Phys. Rev. A 98, 022513 (2018).

Levy–Zahariev shift-in-potential procedure

• Let us consider the N-electron KS-DFT energy expression

$$E_0^N = \sum_{i=1}^N \left\langle \varphi_i \middle| -\frac{\nabla^2}{2} + \mathbf{v}_{\rm ne}(\mathbf{r}) \times \middle| \varphi_i \right\rangle + E_{\rm Hxc}[n]$$
$$= \left(\sum_{i=1}^N \varepsilon_i\right) + E_{\rm Hxc}[n] - \int d\mathbf{r} \frac{\delta E_{\rm Hxc}[n]}{\delta n(\mathbf{r})} n(\mathbf{r})$$

• Shift-in-potential approach*:

$$\frac{\delta E_{\text{Hxc}}[n]}{\delta n(\mathbf{r})} \rightarrow \frac{\delta E_{\text{Hxc}}[n]}{\delta n(\mathbf{r})} + \frac{E_{\text{Hxc}}[n] - \int d\mathbf{r} \ \frac{\delta E_{\text{Hxc}}[n]}{\delta n(\mathbf{r})} n(\mathbf{r})}{\int d\mathbf{r} \ n(\mathbf{r})}$$
$$\frac{\varepsilon_i}{\varepsilon_i} \rightarrow \tilde{\varepsilon_i}$$
$$E_0^N = \sum_{i=1}^N \tilde{\varepsilon_i}$$

- Note that this procedure fixes the KS orbital energy levels. They are *not* defined anymore up to a constant.
- Levy and Zahariev have shown that $I^N = -\tilde{\varepsilon}^N_H + \tilde{w}[n](\mathbf{r})|_{r \to +\infty}$ where $\tilde{w}[n](\mathbf{r})$ is the shifted Hxc potential.
- * M. Levy and F. Zahariev, Phys. Rev. Lett. 113, 113002 (2014).

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Two-weight N-centered ensemble DFT

 In order to extract both IP and EA separately, we consider the two-weight N-centered ensemble density,

$$n_{0}^{\xi}(\mathbf{r}) = \xi_{-} n_{0}^{N-1}(\mathbf{r}) + \xi_{+} n_{0}^{N+1}(\mathbf{r}) + \left[1 - \xi_{-} \frac{N-1}{N} - \xi_{+} \frac{N+1}{N}\right] n_{0}^{N}(\mathbf{r})$$

where $\boldsymbol{\xi} \equiv (\xi_{-}, \xi_{+})$ and the following convexity condition is fulfilled,

 $0 \leq (N-1)\xi_{-} + (N+1)\xi_{+} \leq N.$

- Note that the latter density integrates to *N*.
- Extension of DFT to *N*-centered ground-state ensembles:

$$E_{\rm xc}[n] = E_{\rm xc}^{N,\xi=0}[n] \quad \stackrel{\xi_+>0}{\underset{\varepsilon \to \infty}{\overset{\xi_->0}{\overset$$

B. Senjean and E. Fromager, Phys. Rev. A 98, 022513 (2018).

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Two-weight N-centered ensemble DFT

• Choosing $\xi = 0$ (i.e. returning to standard *N*-electron DFT) leads to¹

$$E_0^N = \sum_{i=1}^N \tilde{\varepsilon}_i^{N, \boldsymbol{\xi} = 0} \qquad \text{and} \qquad E_0^{N \pm 1} = \sum_{i=1}^{N \pm 1} \left[\tilde{\varepsilon}_i^{N, \boldsymbol{\xi} = 0} + \frac{1}{N \pm 1} \left. \frac{\partial E_{xc}^{N, \boldsymbol{\xi}} \left[n_0^N \right]}{\partial \xi_{\pm}} \right|_{\boldsymbol{\xi} = 0} \right]$$

where $\{\tilde{\varepsilon}_{i}^{N,\xi=0}\}_{i=1,2,...}$ are the Levy–Zahariev shifted KS orbital energies².

• Interestingly, we recover the Levy-Zahariev expression for the IP,

$$\left| I^{N} = E_{0}^{N-1} - E_{0}^{N} = -\tilde{\varepsilon}_{\mathrm{H}}^{N,\xi=0} + \left. \frac{\partial E_{\mathrm{xc}}^{N,\xi} \left[n_{0}^{N} \right]}{\partial \xi_{-}} \right|_{\xi=0} \right|_{\xi=0}$$

where the shifted Hxc potential at position $r \to \infty$ is now written explicitly as a density-functional derivative discontinuity contribution.

¹ B. Senjean and E. Fromager, Phys. Rev. A 98, 022513 (2018).

² M. Levy and F. Zahariev, Phys. Rev. Lett. 113, 113002 (2014).

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How to use the N-centered paradigm in practice ?

- A finite $(N < +\infty)$ and uniform electron gas^{*} is the perfect toy system for modeling DDs.
- Indeed, in the latter case,

$$n^{\xi}(\mathbf{r}) = \xi_{-} \frac{N-1}{V} + \xi_{+} \frac{N+1}{V} + \left[1 - \xi_{-} \frac{N-1}{N} - \xi_{+} \frac{N+1}{N}\right] \times \frac{N}{V}$$
$$= \frac{N}{V}$$
$$= n^{N}(\mathbf{r}) \leftarrow \boldsymbol{\xi}\text{-independent!}$$

• We should be able to extract from this model generalized LDA functionals for ensembles*:

$$\begin{array}{ll} \frac{N}{V} &=& n(\mathbf{r}) \\ \\ N &\leftrightarrow & \alpha(\mathbf{r}) = \frac{1}{C_{\mathrm{F}} n^{5/3}(\mathbf{r})} \left(\tau(\mathbf{r}) - \frac{|\boldsymbol{\nabla} n(\mathbf{r})|^2}{4n(\mathbf{r})} \right), \end{array}$$

where $\tau(\mathbf{r})$ is the kinetic energy density and $C_{\rm F} = \frac{3}{5} (6\pi^2)^{2/3}$.

- Those will exhibit DDs through their ξ dependence.
- *P.-F. Loos, J. Chem. Phys. 146, 114108 (2017).

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1D finite uniform electron gas: the ringium model



P.-F. Loos, C. J. Ball, and P. M. W. Gill, J. Chem. Phys. 140, 18A524 (2014).

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N-electron Hamiltonian:

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1D finite uniform electron gas: the ringium model

• Kohn–Sham orbitals:
$$\Phi_m(\varphi) = \frac{1}{\sqrt{2\pi R}} \exp(im\varphi)$$
 with $m = 0, \pm 1, \pm 2, ...$

• Uniform electron density:
$$n^N = \sum_{m=1}^N |\Phi_m(\varphi)|^2 = \frac{N}{2\pi R}.$$

Accurate Hxc energies can be computed for such a model (generalized LDA functional).

• Application to *non-uniform* 1D systems:

(i) electrons in a box: *boxium* model.

- So far, pure ground-state energies only have been computed*.
- The *extension* of the latter work to *N*-centered ensembles is in progress.

*P.-F. Loos, C. J. Ball, and P. M. W. Gill, J. Chem. Phys. 140, 18A524 (2014).

Individual energy levels in ensemble DFT

• For a given (non-degenerate) *M*-state canonical ensemble, the ensemble energy reads

$$E^{\mathbf{w}} = \left(1 - \sum_{i=1}^{M-1} w_i\right) E_0^{\mathbf{N}} + \sum_{i=1}^{M-1} w_i E_i^{\mathbf{N}}.$$

Similarly, the ensemble density reads

$$n^{w}(\mathbf{r}) = \left(1 - \sum_{i=1}^{M-1} w_i\right) n_0^{N}(\mathbf{r}) + \sum_{i=1}^{M-1} w_i n_i^{N}(\mathbf{r}).$$

• Exact energy expression for state *j*:

$$E_j = E^{w} + \sum_{i=1}^{M-1} \left(\delta_{ij} - w_i
ight) rac{\partial E^{w}}{\partial w_i}$$
 , thus lead

$$E_{j} = \tilde{\mathcal{E}}_{j}^{\text{KS}, \mathbf{w}} + \sum_{i=1}^{M-1} \left(\delta_{ij} - w_{i} \right) \left. \frac{\partial E_{\text{xc}}^{\mathbf{w}}[n]}{\partial w_{i}} \right|_{n=n^{\mathbf{w}}}$$

ling to¹

Important consequences: •

(1) excited-state molecular gradients.

(2) Gould-Pittalis-like² decomposition $E_{xc}^{w}[n] = \sum_{i=0}^{M-1} w_i E_{xc}^{i}[n]$ of the ens. xc energy.

² T. Gould and S. Pittalis, arXiv:1808.04994 (2018).

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¹ K. Deur and E. Fromager, to be submitted (2018).

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• École doctorale des sciences chimiques de Strasbourg





*https://quantique.u-strasbg.fr/MCFUNEX

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